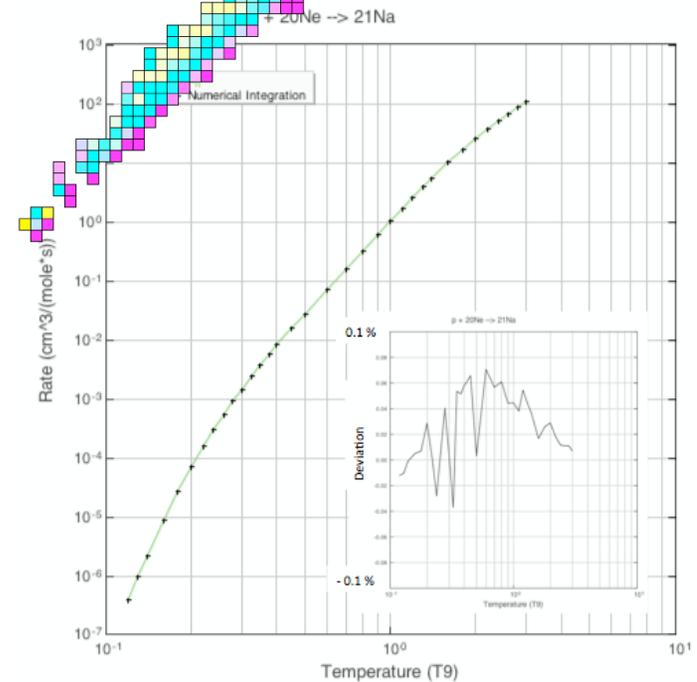
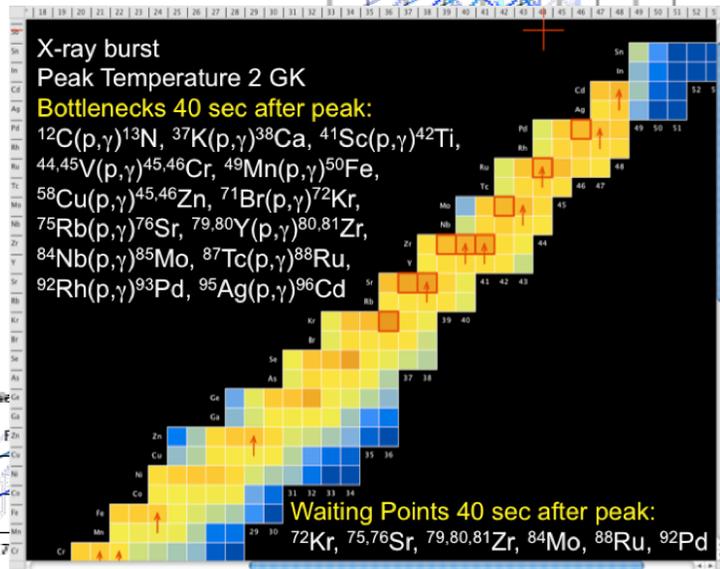
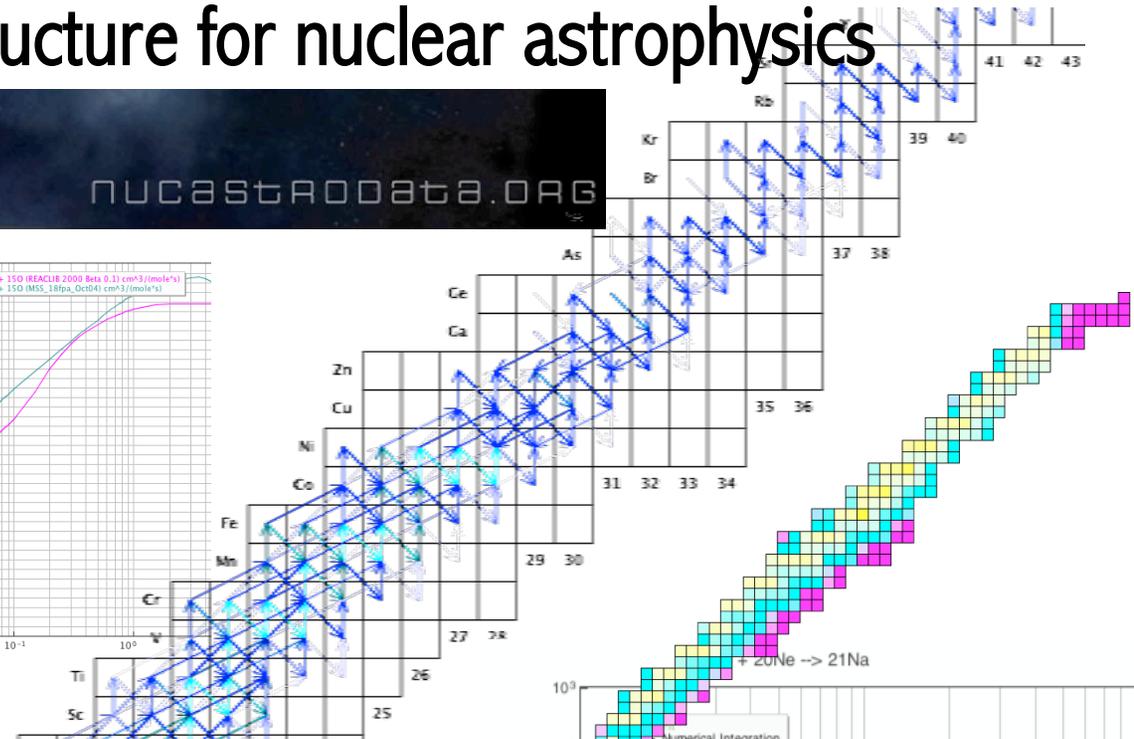
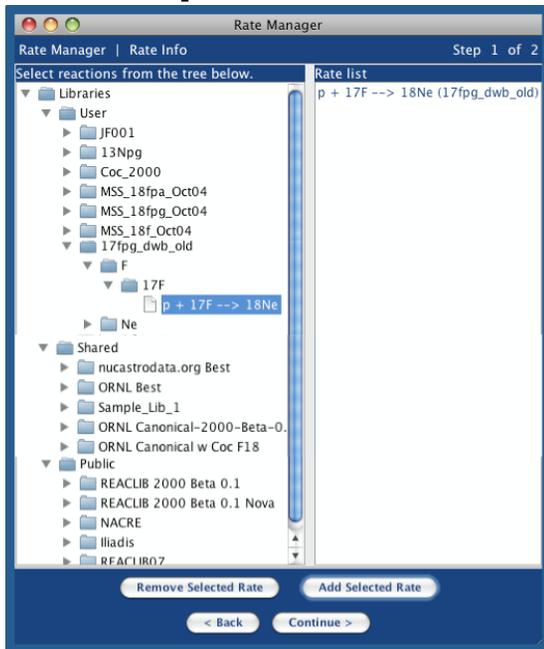


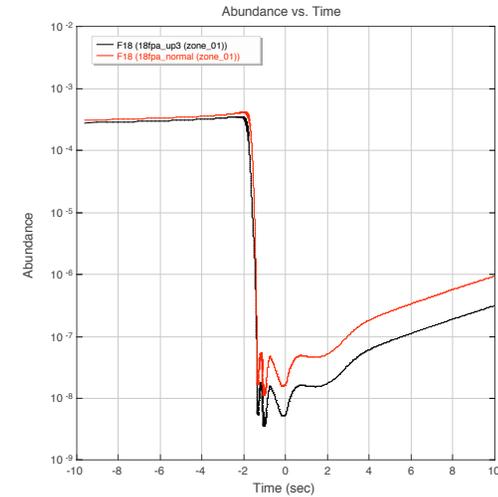
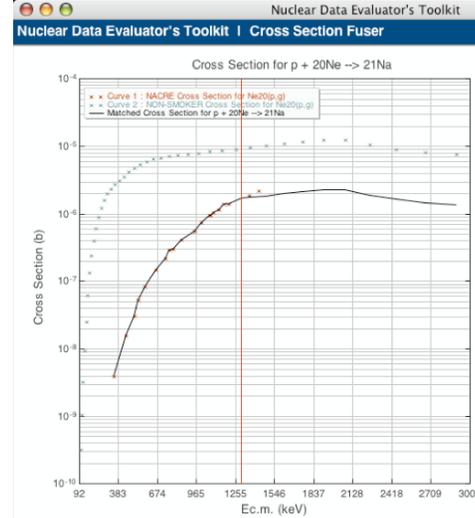
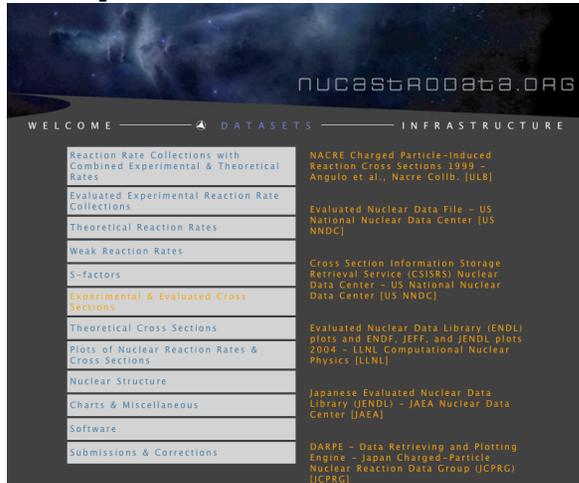
# computational infrastructure for nuclear astrophysics



Michael Smith, Physics Division, Oak Ridge National Lab

[coordinator@nucastrodata.org](mailto:coordinator@nucastrodata.org)

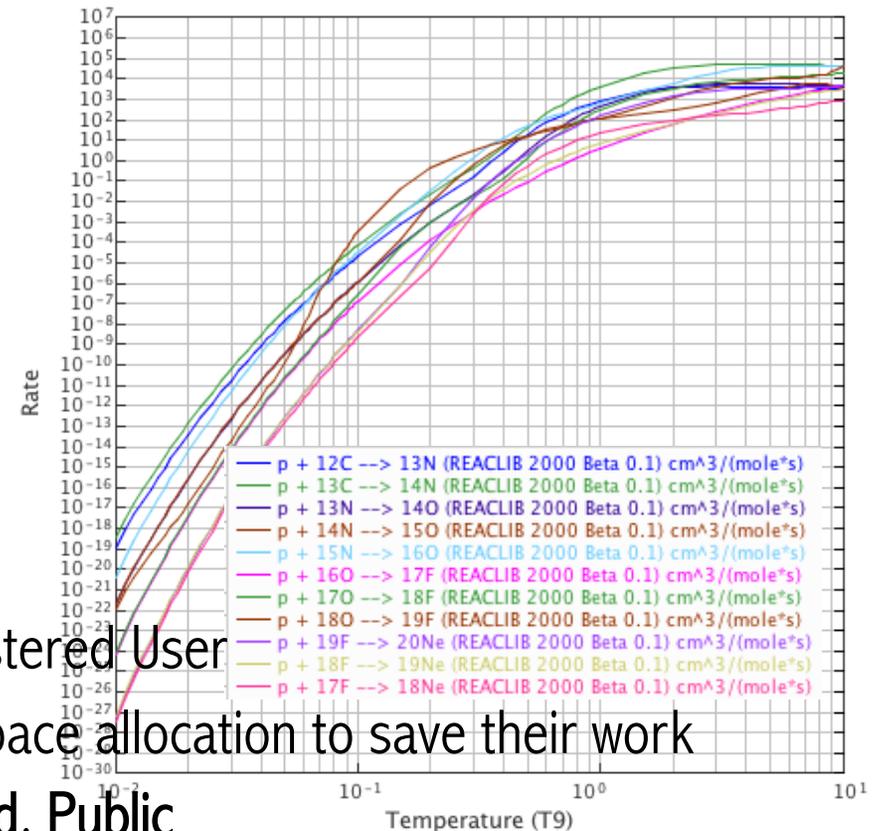
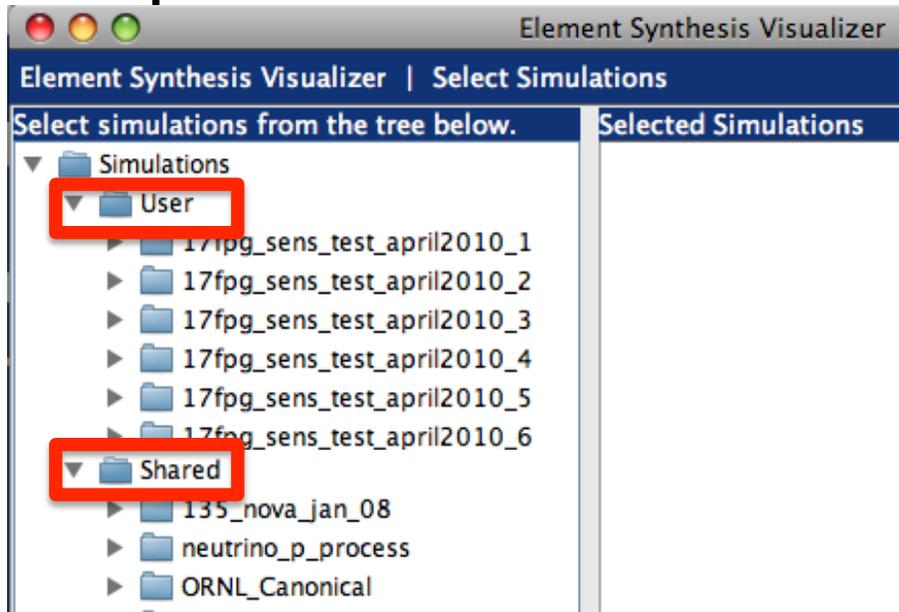
# computational infrastructure for nuclear astrophysics



## system overview

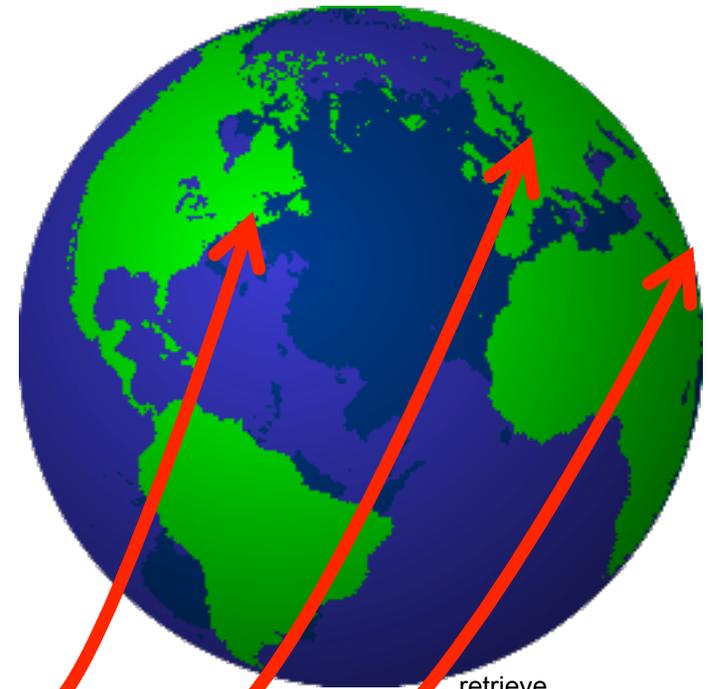
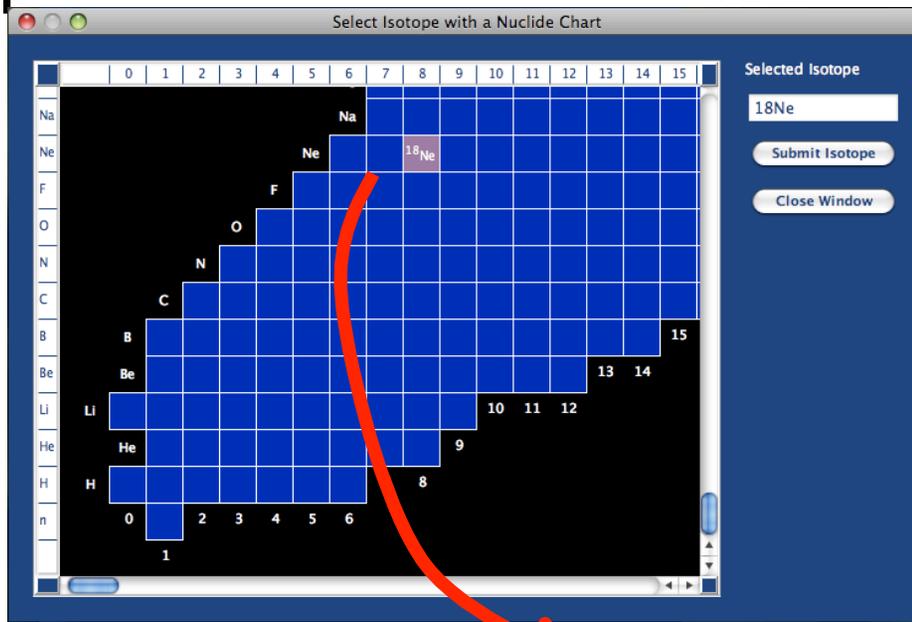
- work with nuclear information, reaction rates, & simulations
- operates “in the cloud” — you need an internet connection
- platform independent java application
- super easy to follow graphical user interface —  
    makes complex manipulations and tasks easy !
- online since 2004; Users from 126 institutions in 29 countries
- new features are always being added at request of Users
- download java program for free at [nuastrodata.org](http://nuastrodata.org)
- contact coordinator @ [nuastrodata.org](mailto:nuastrodata.org) for help

# computational infrastructure for nuclear astrophysics



- log in as a Guest or as Registered User
- Registered Users get disk space allocation to save their work
- 3 data volumes: **User**, **Shared**, **Public**
  - **User** space is for your own rates / libraries / simulations
  - **Shared** space allows you to share this with any other User
  - **Public** space contains published rates / libraries / simulations
- enables easy sharing of large datasets between Users forming an online community

# computational infrastructure for nuclear astrophysics



get data

retrieve info from international databases

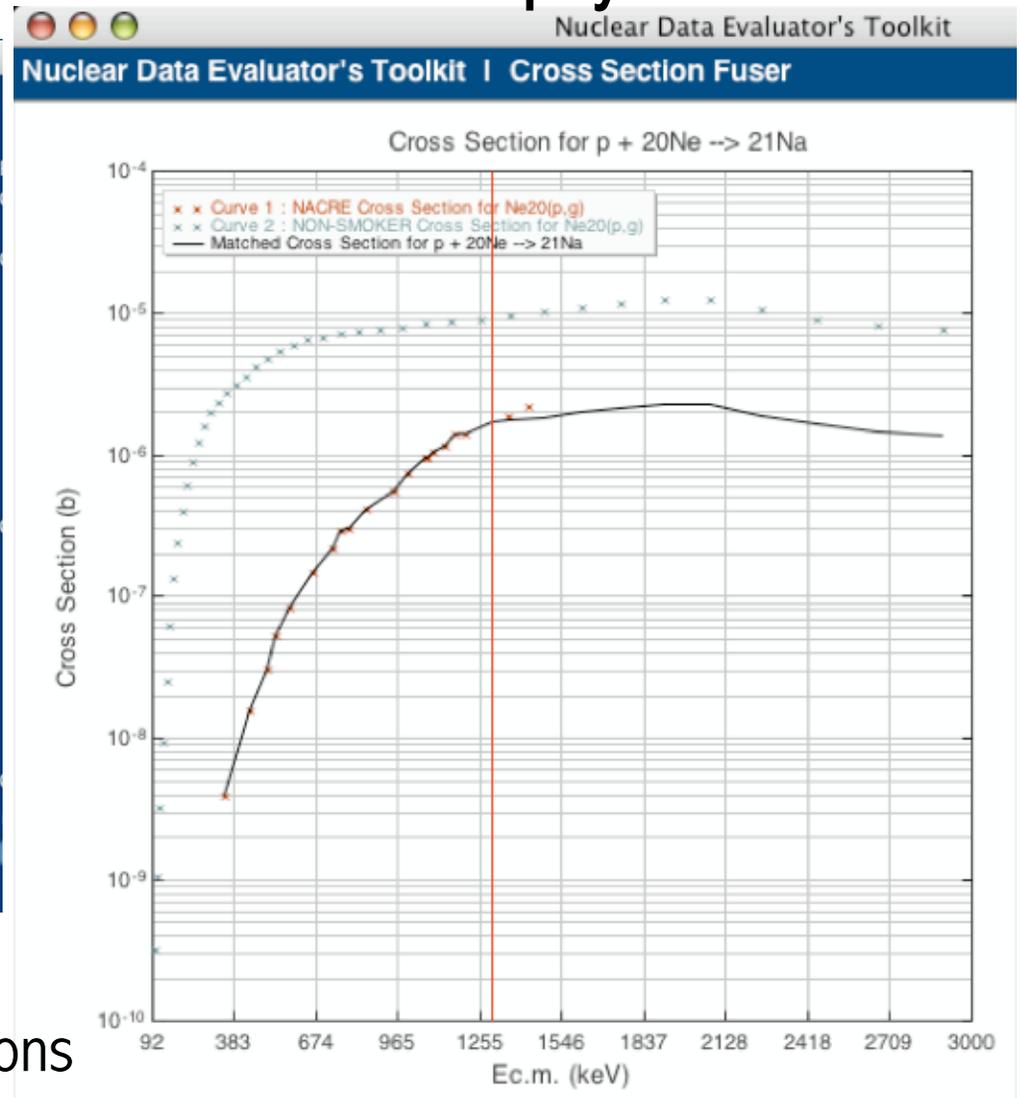
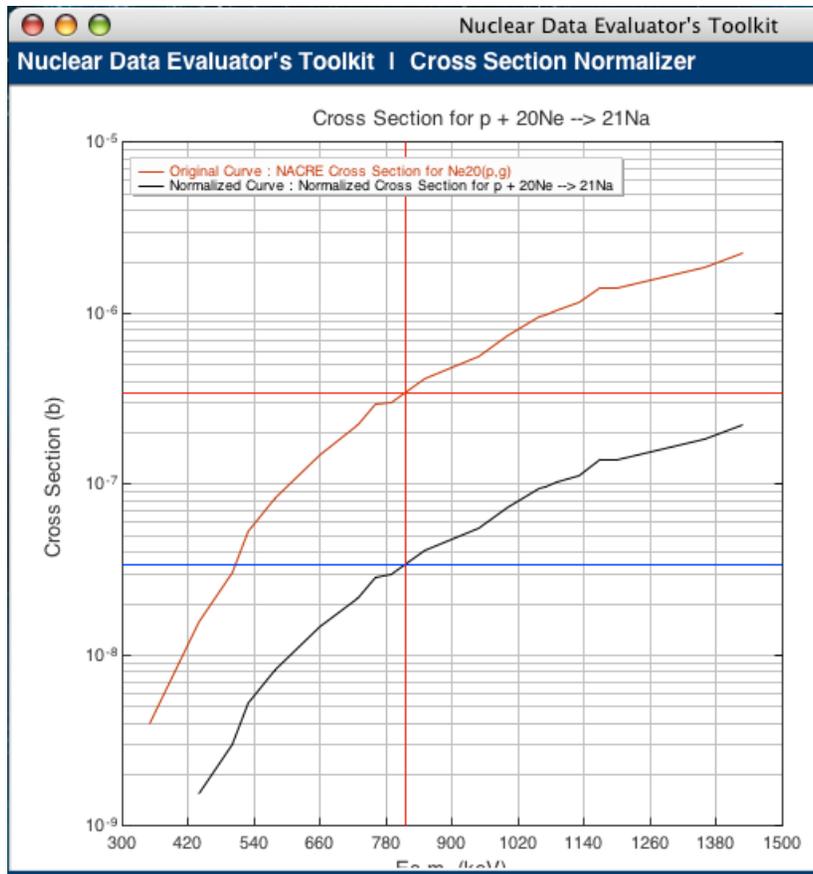


system overview



- new File Repository enables Users to share other types of files
- new “Data Harvester” feature will collect information on chosen nuclei from a number of standard international databases
- extensive workflow tools to streamline reaction evaluations are also online

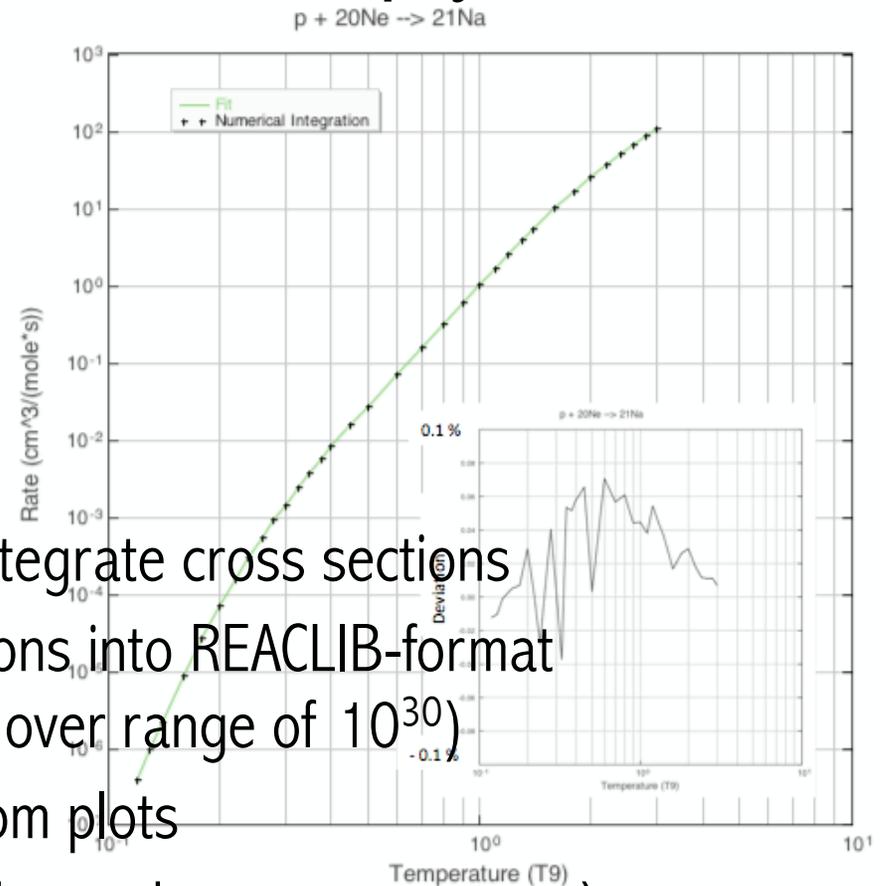
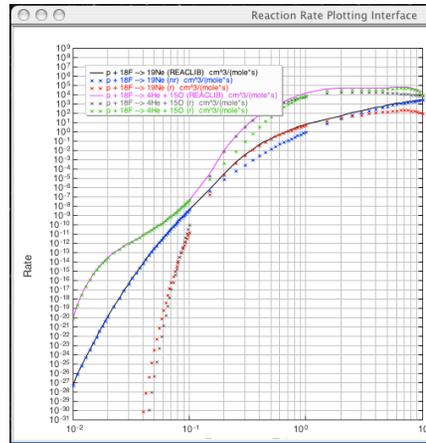
# computational infrastructure for nuclear astrophysics



work with nuclear information

- import cross sections
- gain match, normalize, linearly extrapolate cross sections
- extrapolate experimental cross section with theory

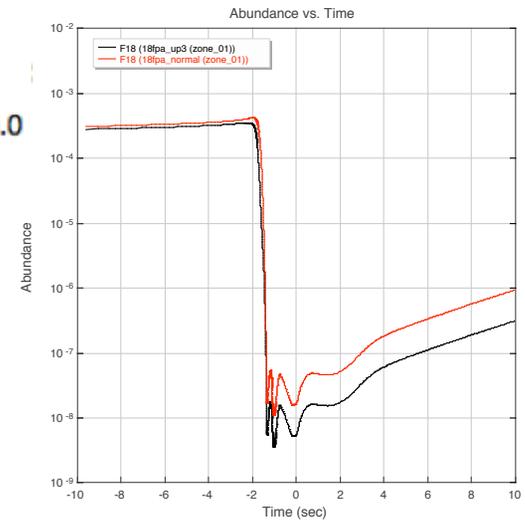
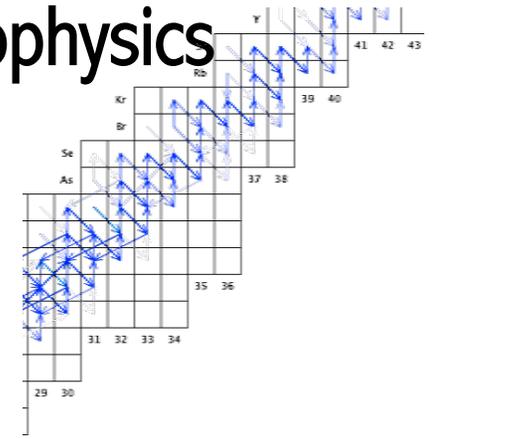
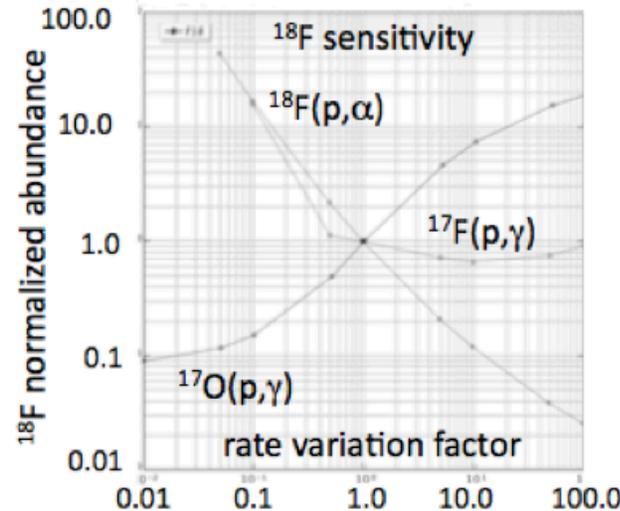
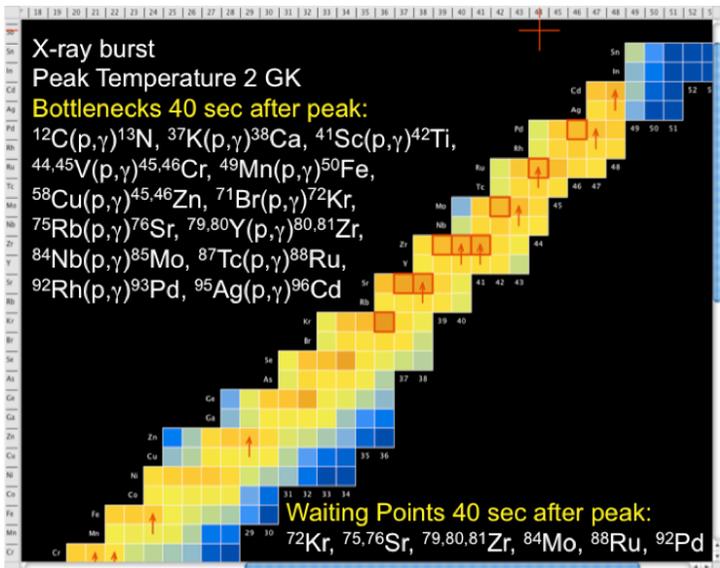
# computational infrastructure for nuclear astrophysics



work with reaction rates

- accurately numerically integrate cross sections
- accurately fit cross sections into REACLIB-format parameterization (2% over range of  $10^{30}$ )
- visualize rates with custom plots
- modify rates (scale up, down, change parameters)
- combine rates into rate libraries for input into simulations
- merge libraries together with custom rules
- save libraries and share with colleagues
- JINA REACLIB v2.0 is the default rate library

# computational infrastructure for nuclear astrophysics

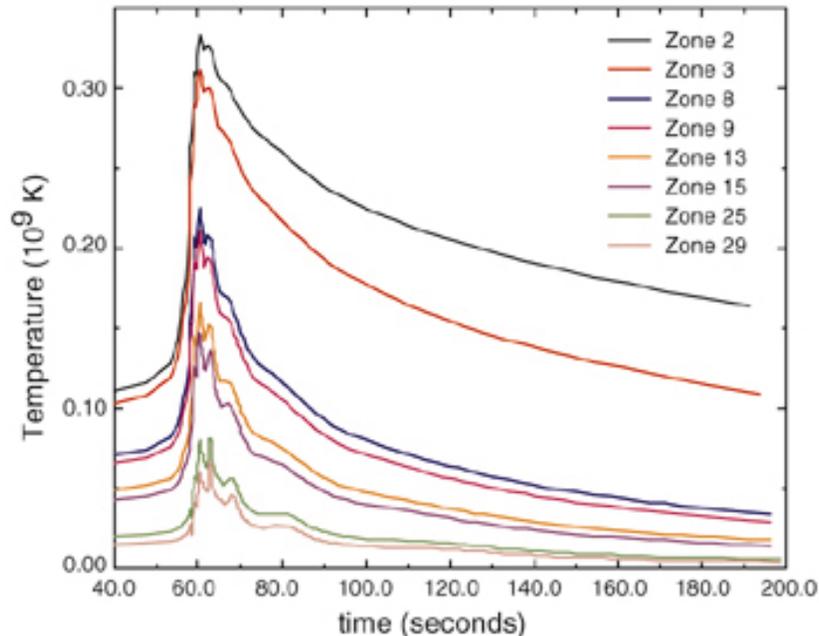


work with simulations

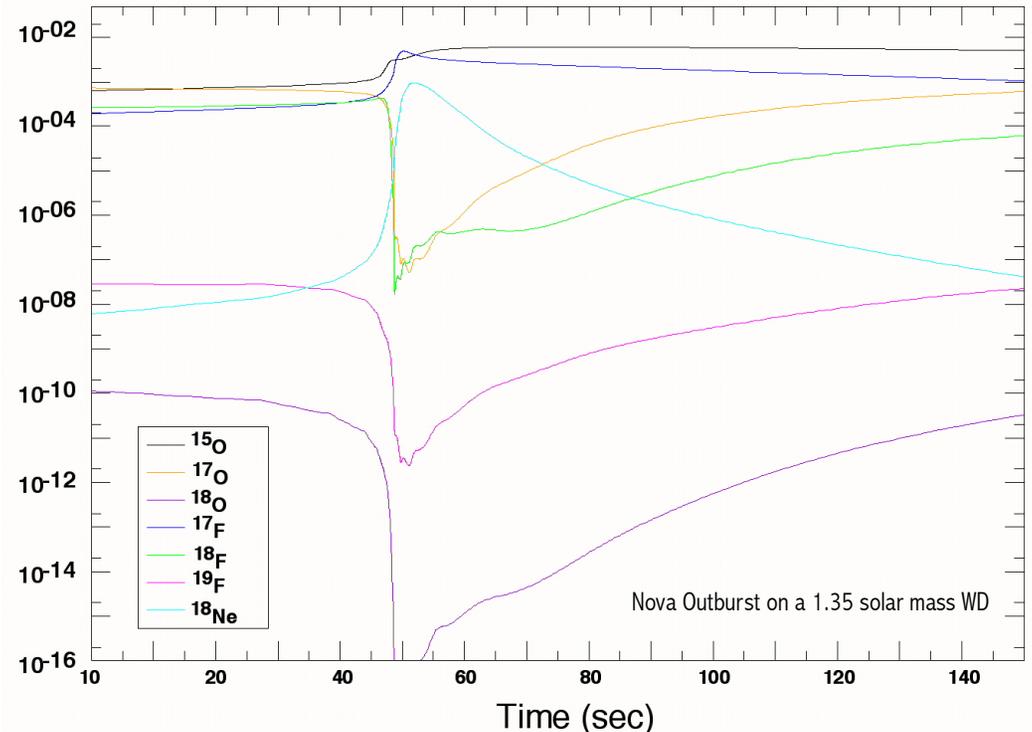
- easily set up and run simulations
- visualize simulations with 1D, 2D plots and animations
- quickly compare simulations with different input
- run **automated sensitivity studies** (changing input rates)
- analyze simulation results – find **bottlenecks, waiting points**

# computational infrastructure for nuclear astrophysics

Temperature profile for different spatial zones



Abundances vs. Time



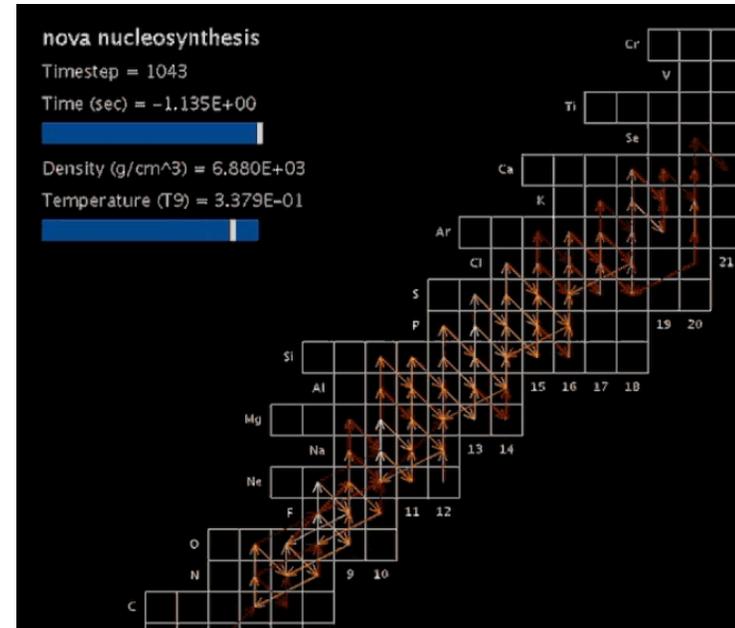
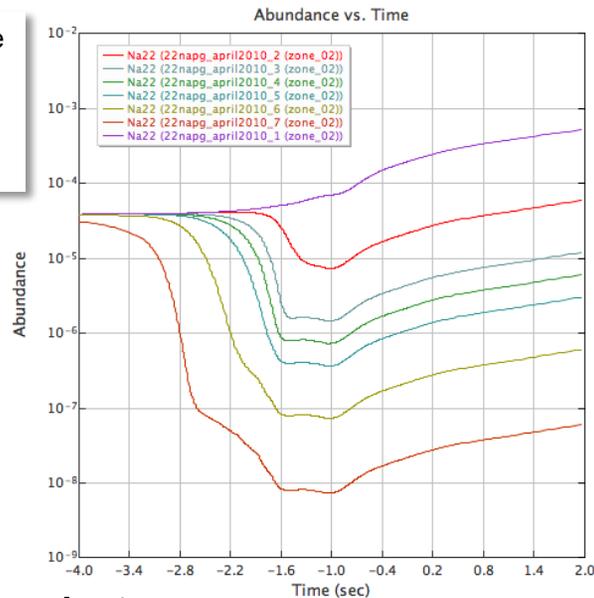
work with simulations

- post-processing code XNET from Raph Hix [ORNL / UTK]
- simulation types: novae, X-ray bursts, solar, CNO, Hot CNO
- single zone and multi-zone simulations
- some simulations with tracer particle temperature/density profiles
- coming soon (1 week): **core collapse supernova r-process**

# computational infrastructure for nuclear astrophysics

Abundance Plotting Interface | Plot Abundance or Ratio of Abundance

$^{22}\text{Na}$  abundance vs. time in Nova outburst with different  $^{22}\text{Na}(p,\gamma)$  rates

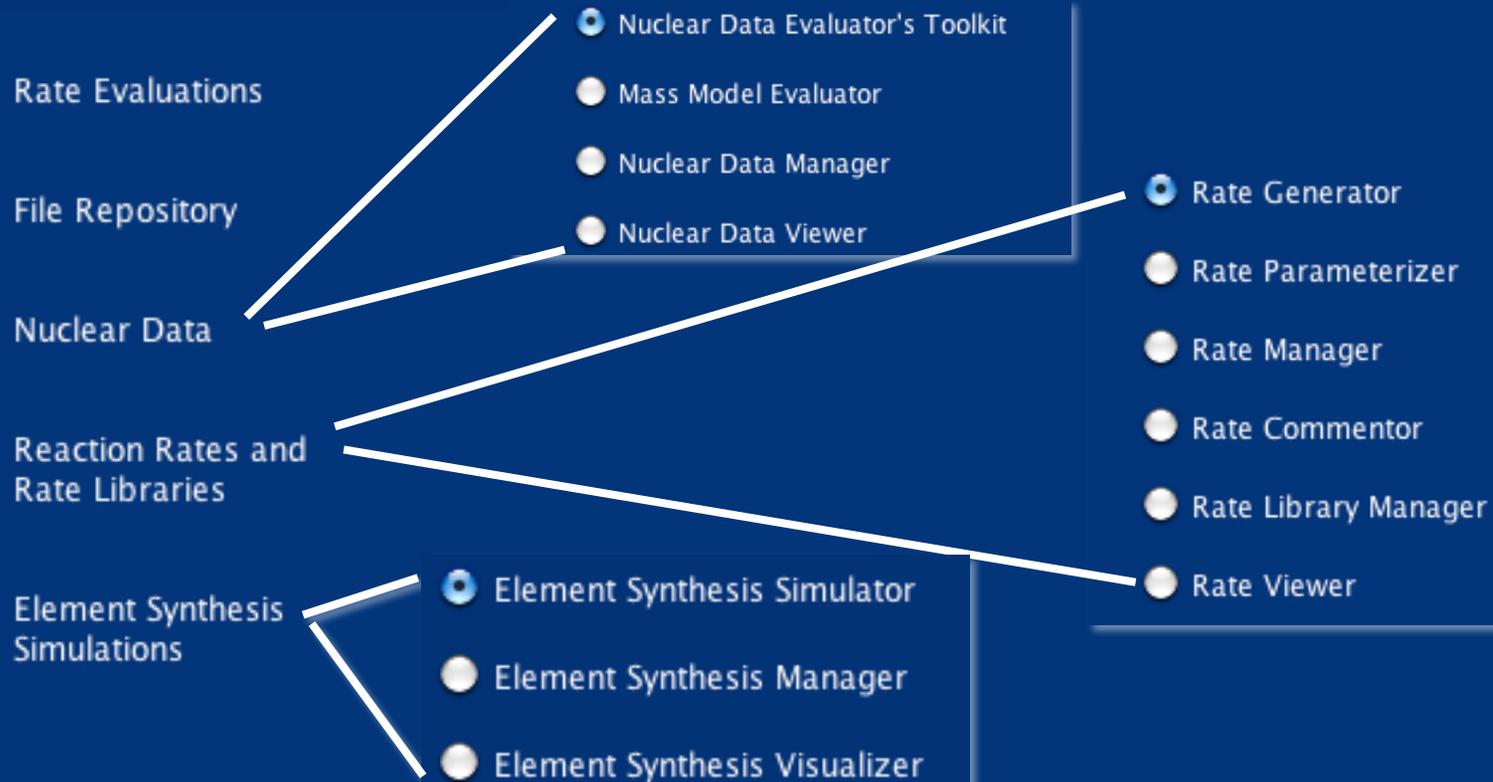


work with simulations

- customized visualization of simulations
- simulations can be saved & shared with colleagues without email or ftp
- you can import your own simulations & use our viz tools
- animation output rendered on our server to your parameters and available for you to download

# computational infrastructure for nuclear astrophysics

## Computational Infrastructure for Nuclear Astrophysics



# sample investigations

## before & after study: what was the impact of my measurement ?

- assume you just measured a new rate of the  $^{18}\text{F}(p,\alpha)^{15}\text{O}$  reaction
- rate is 10 times lower than previous “best” rate
- how does this impact predictions of  $^{18}\text{F}$  production in novae ?

## approach

- modify existing  $^{18}\text{F}(p,\alpha)^{15}\text{O}$  rate appropriately, save into new library
- combine with JINA REACLIB v2.0 into full library for a simulation
- choose a novae simulation, run with old rate & new rate
- compare final abundances in the two simulations
- draw your conclusion on the impact of the measurement
- variations: more nova models, different zones or all zones ...

# sample investigations

## sensitivity study: how does one abundance depend on a rate ?

- you just read an article on  $^{22}\text{Na}$  observations in nova ejecta
- you wonder if measurements of  $^{21}\text{Na}(p,\gamma)^{22}\text{Mg}$  and  $^{21}\text{Ne}(p,\gamma)^{22}\text{Na}$  reaction would help clarify the nuclear uncertainties in  $^{22}\text{Na}$  abundance predictions
- what are the sensitivities of these predictions on these rates ?

## approach

- choose a reference rate library and a novae simulation
- choose a set of variations of these the reactions of interest (e.g., 0.01, 0.1, 1.0, 10, 100 ...)
- use automated sensitivity study tool to vary rate and run simulations
- use tools to plot out abundance vs. rate, look for strong correlations
- variations: more nova models, different rate variations, change other rates on neighboring nuclei ...

# sample investigations

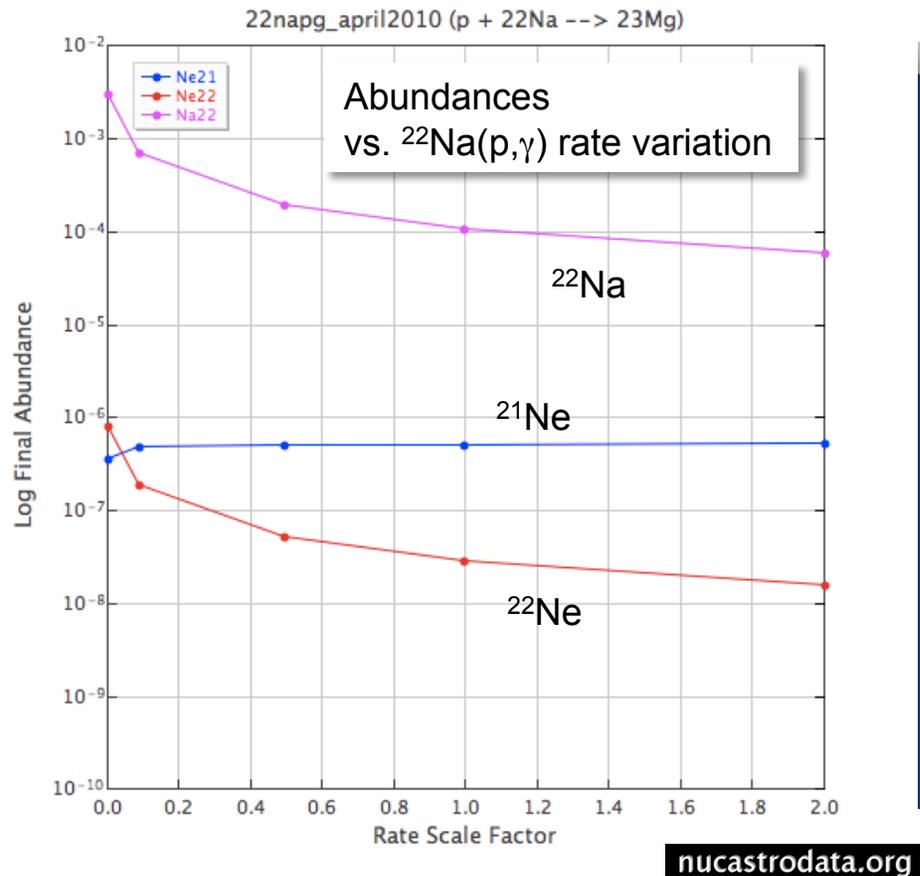
## uncertainty analysis: what is the uncertainty of this predicted abundance ?

- you have just made a new measurement of the  $^{14}\text{O}(\alpha,p)^{17}\text{F}$  reaction with an uncertainty that ranges from a factor of 2 higher to a factor of 20 lower
- what are the implications of this uncertainty for final abundance predictions in novae ?

## approach

- modify  $^{14}\text{O}(\alpha,p)^{17}\text{F}$  rate to have “max” and “min” values
- save each into library, merge with JINA REACLIB v2.0 library
- choose a novae simulation, run with reference, max, & min rates
- compare final abundances, the spread gives the uncertainty
- draw your conclusion on the impact of the uncertainty on  $^{16}\text{O} / ^{18}\text{O}$  and  $^{17}\text{O}/^{18}\text{O}$  abundance ratios, on  $^{18}\text{F}$  and  $^{22}\text{Na}$  abundances ...
- variations: examine different abundances; different nova models ...

# computational infrastructure for nuclear astrophysics



Bottleneck Reaction Finder | Results Step 3 of 3

Below is a list of reactions that are bottlenecks for the synthesis of nuclei with masses greater than or equal to the Bottleneck Mass. Click *Submit Bottleneck Reactions* to visualize these results with the Animator. Click *Close Bottleneck Reaction Finder* to close the Bottleneck Reaction Finder and not submit the results. Check *View Detailed Report* to view the user input as well as the output of the Finder.

Bottleneck Reaction Finder Report:  View Detailed Report

81	$80\text{Y} \rightarrow 81\text{Zr}$	Major Bottleneck
85	$84\text{Nb} \rightarrow 85\text{Mo}$	Major Bottleneck
88	$87\text{Tc} \rightarrow 88\text{Ru}$	Major Bottleneck

Save Copy Print

Submit Bottleneck Reactions

< Back Close Bottleneck Reaction Finder

- download java program for free at [nucastrodata.org](http://nucastrodata.org)
- contact coordinator @ [nucastrodata.org](http://nucastrodata.org) for help
- suggest new features for nuclear data / rates / libraries / simulations / visualization / analysis and we will work with you !